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Lecture Notes on Extra Dimensions

- Kaluza-Klein compactification : one extra dimension on a circle
- Generalizations, orbifolds, symmetry breaking by compactification
- Branes in extra dimensions

**Kaluza-Klein compactification :
one extra dimension on a circle**

X-dimensions

unobservable $\rightarrow$ finite size
 detection : KK spectrum

ex: 1 extra dim on a circle $y \equiv y + 2\pi R$

$$\Rightarrow \phi(x^\mu, y) = \sum_{n \in \mathbb{Z}} \phi_n(x) e^{i y n / R}$$

$$\square_5 \phi = M^2 \phi \quad \text{signature : } (- + + \dots +)$$

$\uparrow$ massive scalar field in 5d

$$\square_5 = \square_4 + \partial_y^2$$

$$\Rightarrow \left(\square - \frac{n^2}{R^2} \right) \phi_n = M^2 \phi_n \quad \Rightarrow \quad \square \phi_n = \left(M^2 + \frac{n^2}{R^2} \right) \phi_n$$

$\uparrow$
 KK spectrum

Effective action : $S = \frac{1}{2} \int d^4 x \int_0^{2\pi R} dy \left\{ -(\partial_\mu \phi)^2 + M^2 \phi^2 \right\}$

$$\int dy e^{i(m+n)y} = 2\pi R \delta_{m+n,0}$$

$$\begin{aligned} \Rightarrow S &= \sum_n \frac{1}{2} \int d^4 x \left\{ 2\pi R \left\{ -(\partial_\mu \phi_n)^2 - \underbrace{\frac{n^2}{R^2} \phi_n^2}_{\frac{n^2}{R^2} \phi_n^2} + M^2 \phi_n^2 \right\} \right\} \\ &= \frac{2\pi R}{2} \int d^4 x \sum_n \left\{ -(\partial_\mu \phi_n)^2 + \left(M^2 + \frac{n^2}{R^2} \right) \phi_n^2 \right\} \end{aligned}$$

normalized field: $\tilde{\phi}_n = \frac{\phi_n}{\sqrt{2\pi R}}$

Unification of forces from 5d (KK idea)

1) Gauge field $\Rightarrow$ gauge + scalar

$$M = (\mu, 4) \longleftrightarrow (x^\mu, y) \quad y \in S^1$$

$$A_M = \begin{pmatrix} A_\mu(x, y) \\ \phi(x, y) \end{pmatrix} \quad \phi \equiv A_4$$

y-dependence: expand in series or before

$$F_{MN}^2 = F_{\mu\nu}^2 + 2 F_{\mu 4}^2$$

$$F_{\mu 4} = \partial_\mu A_4 - \partial_4 A_\mu = \partial_\mu \phi - \partial_y A_\mu$$

5d gauge invariance: $\delta A_M = \partial_M \Lambda(x, y)$

$$\Rightarrow \delta A_\mu = \partial_\mu \Lambda, \quad \delta \phi = \partial_y \Lambda$$

$$\Rightarrow \delta A_\mu^n = \partial_\mu \Lambda_n, \quad \delta \phi_n = i \frac{n}{R} \Lambda_n$$

• $n=0 \Rightarrow$ usual 4d gauge invariance

• $n \neq 0 \Rightarrow$ gauge choice: $\phi_n = 0$ (A_μ^n massive)
longitudinal polarization

$$\Rightarrow F_{\mu 4}^n = \begin{cases} n \neq 0 & i \frac{n}{R} A_\mu^n \\ n=0 & \partial_\mu \phi_0 \end{cases}$$

$$\Rightarrow \int_{2\pi R} dy \frac{1}{4} F_{MN}^2 = \frac{1}{4} F_{\mu\nu}^{(n)2} + \frac{1}{2} (\partial_\mu \phi_0)^2 - \frac{1}{2} \frac{n^2}{R^2} A_\mu^{(n)2}$$

$\Rightarrow$ 4d: U(1) gauge field + ^{massless} scalar + massive KK-vectors $n = \pm 1, \pm 2, \dots$
with masses $^2 \left(\frac{n}{R}\right)^2$

2) gravity 5d $\rightarrow$ gravity 4d + $U(1)$ + scalar (radion)

$$G_{MN} = \begin{pmatrix} G_{\mu\nu} & G_{\mu 4} \\ G_{\mu 4} & G_{44} \end{pmatrix} (x, y)$$

$$\delta G_{MN} = G_{MP} \partial_N \xi^P + G_{NP} \partial_M \xi^P + \xi^P \partial_P G_{MN} \quad \xi^M = \begin{pmatrix} \xi^\mu \\ \xi^4 \end{pmatrix}$$

$$\Rightarrow \begin{cases} \delta G_{44} = 2 G_{4\mu} \partial_4 \xi^\mu + 2 G_{44} \partial_4 \xi^4 + \xi^\mu \partial_\mu G_{44} + \xi^4 \partial_4 G_{44} \\ \delta G_{\mu 4} = G_{\mu\nu} \partial_4 \xi^\nu + G_{\mu 4} \partial_4 \xi^4 + G_{4\nu} \partial_\mu \xi^\nu + G_{44} \partial_\mu \xi^4 \\ \quad + \xi^\nu \partial_\nu G_{\mu 4} + \xi^4 \partial_4 G_{\mu 4} \end{cases}$$

$$\delta G_{44} : \xi^\mu = 0 \Rightarrow \delta G_{44} = 2 G_{44} \partial_4 \xi^4 + \xi^4 \partial_4 G_{44}$$

$$\begin{aligned} \delta G_{44}^n &= \sum_m \left\{ 2 G_{44}^{n-m} \frac{i m}{R} \xi_m^4 + \frac{i(n-m)}{R} \xi_m^4 G_{44}^{n-m} \right\} \\ &= \sum_m \frac{i(n+m)}{R} \xi_m^4 G_{44}^{n-m} \end{aligned}$$

$$G_{44} = e^{2\phi} \Rightarrow \delta \phi = \partial_y \xi^4 + \xi^4 \partial_y \phi$$

$$\Rightarrow \delta \phi_n = \frac{i n}{R} \xi_n^4 + \sum_m \frac{i m}{R} \phi_m \xi_{n-m}^4$$

non-trivial ξ_n^4 , $n \neq 0$ gauge choice : $\phi_n = 0$ (or $G_{44}^n = 0$)

$$\delta G_{\mu 4} : \xi^4 = 0 \Rightarrow \delta G_{\mu 4} = G_{\mu\nu} \partial_4 \xi^\nu + G_{4\nu} \partial_\mu \xi^\nu + \xi^\nu \partial_\nu G_{\mu 4}$$

non-trivial ξ_n^μ , $n \neq 0$ gauge choice : $G_{\mu 4}^n = 0$

$\Rightarrow G_{\mu\nu}^n$: massive spin-2 (5 helicities : $\pm 2, \pm 1, 0$)

$\begin{matrix} \nearrow & & \searrow \\ G_{\mu 4}^n & & G_{44}^n \end{matrix}$

with gauge choice $\Rightarrow$

$$G_{\mu\nu}(x,y) = e^{2\phi(x)} \begin{pmatrix} \hat{G}_{\mu\nu}(x,y) & A_\mu(x) \\ A_\nu(x) & 1 \end{pmatrix} \quad \begin{aligned} G_{44} &= e^{2\phi} \\ G_{\mu 4} &= e^{2\phi} A_\mu \end{aligned}$$

leftover gauge freedom $\xi^\mu(x), \xi^4(x) \quad \delta x^\mu = -\xi^\mu, \quad \delta y = -\xi^4$

$$\delta \hat{G}_{\mu\nu} = \delta (e^{-2\phi} G_{\mu\nu}) =$$

$$\delta G_{44} = \xi^\mu \partial_\mu G_{44} \Rightarrow \boxed{\delta \phi = \xi^\mu \partial_\mu \phi}$$

$$\delta G_{\mu 4} = G_{4\nu} \partial_\mu \xi^\nu + G_{44} \partial_\mu \xi^4 + \xi^\nu \partial_\nu G_{\mu 4} \Rightarrow$$

$$\begin{aligned} \delta A_\mu &= \delta (e^{-2\phi} G_{\mu 4}) = e^{-2\phi} \delta G_{\mu 4} - 2 e^{-2\phi} G_{\mu 4} \delta \phi \\ &= A_\nu \partial_\mu \xi^\nu + \partial_\mu \xi^4 + \underbrace{e^{-2\phi} \xi^\nu \partial_\nu (e^{2\phi} A_\mu)}_{\xi^\nu \partial_\nu A_\mu + 2 A_\mu \xi^\nu \partial_\nu \phi} - 2 A_\mu \xi^\nu \partial_\nu \phi \\ &\Rightarrow \boxed{\delta A_\mu = A_\nu \partial_\mu \xi^\nu + \xi^\nu \partial_\nu A_\mu + \partial_\mu \xi^4} \end{aligned}$$

$$\delta G_{\mu\nu} = G_{\mu\rho} \partial_\nu \xi^\rho + G_{\nu\rho} \partial_\mu \xi^\rho + \xi^\rho \partial_\rho G_{\mu\nu}$$

$$= G_{(\mu\lambda} \partial_{\nu)} \xi^\lambda + \xi^\lambda \partial_\lambda G_{\mu\nu} + G_{(\mu 4} \partial_{\nu)} \xi^4 + \xi^4 \partial_4 G_{\mu\nu}$$

$$\begin{aligned} \Rightarrow \delta \hat{G}_{\mu\nu} &= e^{-2\phi} \delta G_{\mu\nu} - 2 e^{-2\phi} G_{\mu\nu} \delta \phi \\ &= \hat{G}_{(\mu\lambda} \partial_{\nu)} \xi^\lambda + \xi^\lambda \partial_\lambda \hat{G}_{\mu\nu} + \cancel{2 (\xi^\lambda \partial_\lambda \hat{G}_{\mu\nu})} - \cancel{2 \hat{G}_{\mu\nu} \xi^\lambda \partial_\lambda \phi} \\ &\quad + \hat{A}_{(\mu} \partial_{\nu)} \xi^4 + \xi^4 \partial_4 \hat{G}_{\mu\nu} \end{aligned}$$

$$\Rightarrow \boxed{\delta \hat{G}_{\mu\nu} = \hat{G}_{(\mu\lambda} \partial_{\nu)} \xi^\lambda + \xi^\lambda \partial_\lambda \hat{G}_{\mu\nu} + \hat{A}_{(\mu} \partial_{\nu)} \xi^4 + \xi^4 \partial_4 \hat{G}_{\mu\nu}} \quad ((\mu, \nu) = \mu\nu + \nu\mu)$$

$\Rightarrow \xi^\mu(x)$: 4d diffeomorphisms, $\xi^4(x)$: 4d U(1) gauge transformations
(from 5d diffeos)

$$ds^2 = e^{2\phi} d\hat{s}^2$$

$$\begin{aligned} d\hat{s}^2 &= \hat{G}_{\mu\nu} dx^\mu dx^\nu + 2A_\mu dx^\mu dy + dy dy \\ &= (\hat{G}_{\mu\nu} - A_\mu A_\nu) dx^\mu dx^\nu + (dy + A_\mu dx^\mu)^2 \end{aligned}$$

$$\delta_4 (dy + A_\mu dx^\mu) = \cancel{\partial_\mu y} dx^\mu + \partial_\mu \xi^4 + \partial_\mu \xi^4 dx^\mu = 0$$

$$\delta_4 \left(\underbrace{\hat{G}_{\mu\nu} - A_\mu A_\nu}_{\hat{g}_{\mu\nu}} \right) = \cancel{A_\mu \partial_\nu \xi^4} + \xi^4 \partial_\mu \hat{G}_{\mu\nu} - \cancel{A_\mu \partial_\nu \xi^4} = \xi^4 \partial_\mu \hat{G}_{\mu\nu} = 0$$

for $\hat{G}_{\mu\nu}$

$$\Rightarrow G_{MN} = e^{2\phi} \begin{pmatrix} \hat{g}_{\mu\nu} + A_\mu A_\nu & A_\mu \\ A_\nu & 1 \end{pmatrix}$$

$$R[e^{2\phi} g_{\mu\nu}] = e^{-2\phi} \left\{ R - 2(D-1) \partial^2 \phi - (D-2)(D-1)(\partial\phi)^2 \right\}$$

$$\sqrt{g}[e^{2\phi} g_{\mu\nu}] = e^{D\phi} \sqrt{g}$$

$$\Rightarrow \sqrt{g} R^{(5)}[e^{2\phi} \hat{G}_{\mu\nu}] = e^{3\phi} \left\{ R^{(5)} - 8 \partial_\mu^2 \phi - 12 (\partial_\mu \phi)^2 \right\}$$

$$\left[\begin{aligned} R^\lambda_{\mu\nu\rho} &= \partial_\rho \Gamma^\lambda_{\mu\nu} - \partial_\nu \Gamma^\lambda_{\mu\rho} + \Gamma^\sigma_{\mu\nu} \Gamma^\lambda_{\rho\sigma} - \Gamma^\sigma_{\mu\rho} \Gamma^\lambda_{\nu\sigma} & R_{\mu\rho} &= R^\lambda_{\mu\lambda\rho} \\ \Gamma^{\sigma\tau}_{\lambda\mu} &= \frac{1}{2} g^{\sigma\nu} (\partial_\lambda g_{\mu\nu} + \partial_\mu g_{\lambda\nu} - \partial_\nu g_{\lambda\mu}) & R &= g^{\mu\rho} R^\lambda_{\mu\lambda\rho} \end{aligned} \right]$$

$$\hat{G} = \begin{pmatrix} \hat{g}_{ij} + A_i A_j & A_i \\ A_i^T & 1 \end{pmatrix} = \begin{pmatrix} \vec{g}_1 + \vec{A} A_1 & \vec{g}_2 + \vec{A} A_2 & \dots & \vec{g}_n + \vec{A} A_n & \vec{A} \\ A_1 & A_2 & & A_n & 1 \end{pmatrix}$$

$$\cancel{\hat{g}_{ij} + A_i A_j}$$

$$\vec{g}_i = \begin{pmatrix} g_{i1} \\ g_{i2} \\ \vdots \\ g_{in} \end{pmatrix}$$

$$1) \det \begin{pmatrix} \vec{g}_1 + \vec{A} A_1 & \vec{g}_2 + \vec{A} A_2 & \dots & \vec{g}_n + \vec{A} A_n & \vec{A} \\ & A_1 & & A_n & 1 \end{pmatrix} = \det \hat{G}$$

$$= \det \begin{pmatrix} \vec{g}_1 & \dots & \vec{g}_n & \vec{A} \\ 0 & & 0 & 1 \end{pmatrix} = \det \begin{pmatrix} \vec{g} & A \\ & 1 \end{pmatrix} = \det \hat{g}$$

$$2) \det(\hat{g} + A \otimes A^T) = \det \hat{g} + \det(1 + \hat{g}^{-1} A \otimes A^T)$$

$$\det(1 + B \otimes A^T) = \det \left(\vec{e}_1 + A_1 \vec{B}, \vec{e}_2 + A_2 \vec{B}, \dots, \vec{e}_n + A_n \vec{B} \right)$$

$$e_i = \begin{pmatrix} 0 \\ \vdots \\ 1 \\ \vdots \\ 0 \end{pmatrix} \leftarrow i\text{-th line}$$

$$= \det \left(\vec{e}_1 + A_1 \vec{B}, \vec{e}_2 - \frac{A_2}{A_1} \vec{e}_1, \dots, \vec{e}_n - \frac{A_n}{A_1} \vec{e}_1 \right)$$

$$= \begin{vmatrix} 1 + A_1 B_1 & -A_2/A_1 & \dots & -A_n/A_1 \\ A_1 B_2 & 1 & & 0 \\ \vdots & & \ddots & \vdots \\ A_1 B_n & 0 & & 1 \end{vmatrix} = -(-1)^n \frac{A_n}{A_1} \begin{vmatrix} A_1 B_1 & 1 \\ & \ddots \\ A_1 B_n & 1 \end{vmatrix}$$

M_n

$$+ \begin{vmatrix} 1 + A_1 B_1 & \dots & -A_{n-1}/A_1 \\ & \ddots & \\ A_1 B_{n-1} & & 1 \end{vmatrix}$$

M_{n-1}

$$\Rightarrow \det M_n = \det M_{n-1} - (-1)^n \frac{A_n}{A_1} (-1)^{n-1} A_1 B_n \det 1 = \det M_{n-1} + A_n B_n$$

$$= \dots = \vec{A} \cdot \vec{B}$$

$$\Rightarrow \det(1 + \hat{g}^{-1} A \otimes A^T) = \hat{A}^T \hat{g}^{-1} A$$

$$\Rightarrow \det(\hat{g} + A \otimes A^T) = \det \hat{g} + \hat{A}^T \hat{g}^{-1} A$$

$$\Rightarrow \hat{G}^{-1} = \begin{pmatrix} \hat{g}^{-1} & -\hat{g}^{-1} A \\ -\hat{A}^T \hat{g}^{-1} & 1 + \hat{A}^T \hat{g}^{-1} A \end{pmatrix}$$

$$\begin{aligned} * \begin{pmatrix} \hat{g} + A \otimes A^T & A \\ A^T & 1 \end{pmatrix} \begin{pmatrix} \hat{g}^{-1} & -\hat{g}^{-1} A \\ -\hat{A}^T \hat{g}^{-1} & 1 + \hat{A}^T \hat{g}^{-1} A \end{pmatrix} &= \\ = \begin{pmatrix} 1 + A \otimes \hat{A}^T \hat{g}^{-1} A - A \otimes \hat{A}^T \hat{g}^{-1} A & -A - A \hat{A}^T \hat{g}^{-1} A + A + A \hat{A}^T \hat{g}^{-1} A \\ \hat{A}^T \hat{g}^{-1} A - \hat{A}^T \hat{g}^{-1} A & -\hat{A}^T \hat{g}^{-1} A + 1 + \hat{A}^T \hat{g}^{-1} A \end{pmatrix} & \text{OK} \end{aligned}$$

fields: g -independent

$$\begin{aligned} \Gamma_{\mu\nu}^\sigma &= \hat{\Gamma}_{\mu\nu}^\sigma + \frac{1}{2} \hat{G}^{\sigma\tau} \left\{ \partial_\mu (A_\tau A_\nu) + \partial_\nu (A_\tau A_\mu) - \partial_\tau (A_\mu A_\nu) \right\} + \frac{1}{2} \hat{G}^{\sigma 4} (\partial_\mu A_\nu + \partial_\nu A_\mu - \partial_\mu A_\nu) \\ &= \hat{\Gamma}_{\mu\nu}^\sigma - \frac{1}{2} A^\sigma (\partial_\mu A_\nu + \partial_\nu A_\mu) + \frac{1}{2} \hat{g}^{\sigma\tau} \underbrace{(A_\mu \partial_\nu A_\tau + A_\nu \partial_\mu A_\tau + A_\mu \partial_\tau A_\nu + A_\nu \partial_\tau A_\mu - A_\mu \partial_\nu A_\tau - A_\nu \partial_\mu A_\tau)}_{= A_\tau F_{\mu\nu} + A_\mu F_{\tau\nu} + A_\nu F_{\tau\mu}} \\ &\Rightarrow \Gamma_{\mu\nu}^\sigma = \hat{\Gamma}_{\mu\nu}^\sigma + \frac{1}{2} \hat{g}^{\sigma\tau} (A_\tau F_{\mu\nu} + A_\mu F_{\tau\nu} + A_\nu F_{\tau\mu}) \end{aligned}$$

$$\Gamma_{44}^\Sigma = \frac{1}{2} \hat{G}^{\Sigma N} (\partial_4 \hat{G}_{4N} - \partial_N \hat{G}_{44}) = 0$$

$$\Gamma_{4\mu}^\Sigma = \frac{1}{2} \hat{G}^{\Sigma N} (\partial_4 \hat{G}_{\mu N} + \partial_\mu \hat{G}_{4N} - \partial_N \hat{G}_{\mu 4}) = \frac{1}{2} \hat{G}^{\Sigma\nu} F_{\mu\nu} \quad \begin{cases} \Gamma_{4\mu}^\sigma = \frac{1}{2} \hat{g}^{\sigma\nu} F_{\mu\nu} \\ \Gamma_{4\mu}^4 = -\frac{1}{2} A^\nu F_{\mu\nu} \end{cases}$$

$$\begin{aligned} \Gamma_{\mu\nu}^4 &= \frac{1}{2} \hat{G}^{44} (\partial_\mu A_\nu + \partial_\nu A_\mu) + \frac{1}{2} \hat{G}^{4\tau} (\partial_\mu \hat{G}_{\tau\nu} + \partial_\nu \hat{G}_{\tau\mu} - \partial_\tau \hat{G}_{\mu\nu}) \\ &= \frac{1}{2} (1 + A^\tau A_\tau) (\partial_\mu A_\nu + \partial_\nu A_\mu) - \frac{1}{2} A^\tau \hat{\Gamma}_{\mu\nu}^\tau - \frac{1}{2} A^\tau \left\{ \partial_\mu (A_\tau A_\nu) + \partial_\nu (A_\tau A_\mu) - \partial_\tau (A_\mu A_\nu) \right\} \\ &= -\frac{1}{2} A^\tau \hat{\Gamma}_{\mu\nu}^\tau + \frac{1}{2} (\partial_\mu A_\nu + \partial_\nu A_\mu) - \frac{1}{2} A^\tau (A_\tau F_{\mu\nu} + A_\mu F_{\tau\nu} + A_\nu F_{\tau\mu}) \end{aligned}$$

$$\Gamma_{N4}^N = \Gamma_{\mu 4}^{\mu} + \Gamma_{44}^4 = \frac{1}{2} \int \partial_{\sigma} \int^{\mu \sigma \nu} F_{\mu \nu} = 0$$

$$\Gamma_{N5}^N = \Gamma_{\mu \sigma}^{\mu} + \Gamma_{4\sigma}^4 = \cancel{\Gamma_{\nu \sigma}^{\nu}} + \cancel{\frac{1}{2} A^{\nu} F_{\sigma \nu}} - \frac{1}{2} A^{\nu} F_{\sigma \nu} = \Gamma_{\nu \sigma}^{\nu}$$

$$R = \hat{G}^{\mu \rho} \partial_{\rho} \Gamma_{MN}^N - \hat{G}^{\mu \rho} \partial_{\nu} \Gamma_{MR}^{\nu} + \hat{G}^{\mu \rho} \left(\Gamma_{MN}^{\Sigma} \Gamma_{R\Sigma}^N - \Gamma_{MR}^{\Sigma} \Gamma_{N\Sigma}^N \right)$$

$$= \hat{g}^{\mu \rho} \partial_{\rho} \Gamma_{\mu \nu}^{\nu} - \hat{g}^{\mu \rho} \partial_{\nu} \Gamma_{\mu \rho}^{\nu} - 2 \hat{G}^{\mu 4} \partial_{\nu} \Gamma_{\mu 4}^{\nu} - \hat{G}^{\mu \rho} \Gamma_{MR}^{\sigma} \Gamma_{N\sigma}^N + \hat{G}^{\mu \rho} \Gamma_{MN}^{\Sigma} \Gamma_{R\Sigma}^N$$

$$= \hat{R} - \hat{g}^{\mu \lambda} \partial_{\sigma} \left\{ \hat{g}^{\sigma \nu} A_{\lambda} F_{\mu \nu} \right\} + A^{\mu} \partial_{\sigma} \left(\hat{g}^{\sigma \nu} F_{\mu \nu} \right) - 2 \hat{G}^{\mu 4} \Gamma_{\mu 4}^{\sigma} \Gamma_{\nu \sigma}^{\nu} - \hat{g}^{\mu \rho} \Gamma_{\mu \rho}^{\sigma} \Gamma_{\nu \sigma}^{\nu} + \dots$$

$$= \hat{R} - \hat{g}^{\mu \lambda} \hat{g}^{\sigma \nu} (\partial_{\sigma} A_{\lambda}) F_{\mu \nu} + A^{\mu} \hat{g}^{\sigma \lambda} F_{\mu \lambda} \Gamma_{\nu \sigma}^{\nu} - \hat{g}^{\mu \lambda} \hat{g}^{\sigma \rho} A_{\lambda} F_{\mu \rho} \Gamma_{\nu \sigma}^{\nu} + \dots$$

$$= \hat{R} + \frac{1}{2} F^2 + \hat{G}^{\mu \rho} \left\{ \Gamma_{MN}^{\sigma} \Gamma_{R\sigma}^N + \Gamma_{MN}^4 \Gamma_{R4}^N \right\}$$

$$\Gamma_{\mu \nu}^{\sigma} \Gamma_{R\sigma}^{\nu} + 2 \Gamma_{\mu 4}^{\sigma} \Gamma_{R\sigma}^4 + \Gamma_{\mu 4}^4 \Gamma_{R4}^4$$

$$\left\{ \right\} = \hat{G}^{\mu 4} \Gamma_{\mu \nu}^{\sigma} \Gamma_{R\sigma}^{\nu} + \hat{G}^{\mu 4} \left(2 \Gamma_{\mu \nu}^{\sigma} \Gamma_{R\sigma}^4 + 2 \Gamma_{\mu 4}^{\sigma} \Gamma_{R\sigma}^4 \right) + \left(\Gamma_{\mu \nu}^{\sigma} \Gamma_{R\sigma}^{\nu} + 2 \Gamma_{\mu 4}^{\sigma} \Gamma_{R\sigma}^4 + \Gamma_{\mu 4}^4 \Gamma_{R4}^4 \right) \hat{g}^{\mu \rho}$$

$$= (1+A^2) \frac{1}{4} \hat{g}^{\sigma \lambda} F_{\nu \lambda} \hat{g}^{\nu \rho} F_{\sigma \rho} - A^{\lambda} \left(\hat{g}^{\nu \lambda} F_{\sigma \lambda} \Gamma_{\mu \nu}^{\sigma} - \frac{1}{2} \hat{g}^{\sigma \lambda} F_{\mu \lambda} A^{\rho} F_{\sigma \rho} \right) + \Gamma_{\mu \nu}^{\sigma} \Gamma_{R\sigma}^{\nu} \hat{g}^{\mu \rho}$$

$$+ \hat{g}^{\mu \rho} \hat{g}^{\sigma \nu} F_{\mu \nu} \Gamma_{R\sigma}^4 + \frac{1}{4} A^{\nu} F_{\mu \nu} A^{\lambda} F_{\rho \lambda} \hat{g}^{\mu \rho}$$

$$= -(1+A^2) \frac{1}{4} F^2 - \frac{1}{4} (A^{\nu} F_{\mu \nu})^2 - A^{\mu} \Gamma_{\mu \nu}^{\sigma} \hat{g}^{\nu \lambda} F_{\sigma \lambda} - \frac{1}{2} \hat{g}^{\nu \lambda} F_{\sigma \lambda} A^{\mu} \hat{g}^{\sigma \rho} (A_{\mu} F_{\nu \rho} + A_{\nu} F_{\mu \rho})$$

$$+ \hat{g}^{\mu \rho} \Gamma_{R\sigma}^{\nu} \hat{g}^{\sigma \lambda} (A_{\mu} F_{\nu \lambda} + A_{\nu} F_{\mu \lambda}) + \frac{1}{4} \hat{g}^{\mu \rho} \hat{g}^{\sigma \lambda} (A_{\mu} F_{\nu \lambda} + A_{\nu} F_{\mu \lambda}) \hat{g}^{\nu \tau}$$

$$= -\frac{1}{4} (1+A^2) F^2 - \frac{1}{4} (A^{\nu} F_{\mu \nu})^2 + \frac{1}{2} A^2 F^2 + \frac{1}{2} (A^{\nu} F_{\mu \nu})^2 - \frac{1}{4} A^2 F^2 - \frac{1}{2} (A^{\nu} F_{\mu \nu})^2 + \frac{1}{4} (A^{\nu} F_{\mu \nu})^2$$

$$= -\frac{1}{4} F^2$$

$$\Rightarrow \hat{R}^{(5)} = \hat{R}^{(4)} + \frac{1}{4} F^2$$

$$\sqrt{g} R^{(5)} \left[e^{2\phi} \hat{G}_{\mu\nu} \right]_{4d} = e^{3\phi} \sqrt{\hat{G}} \left[\hat{R} + \frac{1}{4} F^2 - 8(\hat{D}\phi)^2 - 12(\hat{D}\phi)^2 \right] \quad (*)$$

$$= e^{3\phi} \sqrt{\hat{G}} \left[\hat{R} + \frac{1}{4} F^2 + 12(\hat{D}\phi)^2 \right]$$

$$\hat{G}_{\mu\nu} = e^{2\omega} g_{\mu\nu} \Rightarrow e^{2\omega} \left[R - 6(D\omega)^2 - 6(D\omega)^2 \right]$$

$$\omega = -\frac{3}{2}\phi \Rightarrow = R - 6 \cdot \frac{9}{4} (D\phi)^2 + 12 (D\phi)^2 + \frac{1}{4} e^{3\phi} F^2$$

$$12 - \frac{27}{2} = -\frac{3}{2} \Rightarrow = R - \frac{3}{2} (D\phi)^2 + \frac{1}{4} e^{3\phi} F^2$$

$$\phi = \frac{1}{\sqrt{3}} \varphi \Rightarrow = R - \frac{1}{2} (D\varphi)^2 + \frac{1}{4} e^{\sqrt{3}\varphi} F^2$$

$$G_{MN} = e^{2/\sqrt{3}\varphi} \begin{pmatrix} e^{-\sqrt{3}\varphi} g_{\mu\nu} + A_\mu A_\nu & A_\mu \\ A_\nu & 1 \end{pmatrix}$$

$$\Rightarrow \frac{1}{2\kappa_5^2} \int dy R^{(5)} = \frac{2\pi R}{2\kappa_5^2} \left\{ R - \frac{1}{2} (D\varphi)^2 + \frac{1}{4} e^{\sqrt{3}\varphi} F^2 + \sum_{n \neq 0} (\text{massive spin-2 fields}) \right\}$$

Kaluza-Klein idea: unify gravity with electromagnetism in 5d

φ : "Radion" excitation

Radon: $\bullet G_{44} = e^{2\phi} \quad (*) \Rightarrow \langle e^{3\phi} \rangle \text{ changes } 2\pi R$

$\Rightarrow \langle e^{\sqrt{3}\varphi} \rangle$ fixes the radius of the extra dimension

EFT: integrate out massive KK modes $\Rightarrow$ EFT for light modes
 $\Rightarrow$ higher dim eff operators

4d $\leftrightarrow$ 5d couplings 1) $U(1): \frac{1}{g_4^2} = \frac{2\pi R}{g_5^2} \quad [g_5^2] = \text{length}$

2) gravity: $M_P^2 = \frac{1}{\kappa_4^2} = \frac{2\pi R}{\kappa_5^2} = 2\pi R M_5^3$

- Force in extra dims

1) Coulomb ~~Q~~ potential

$$\vec{B}=0 \quad \rho = Q \delta^3(x) \quad \vec{\nabla} \cdot \vec{E} = \rho \quad \vec{\nabla} \times \vec{E} = 0 \quad \Rightarrow \quad \vec{E} = \vec{\nabla} \phi, \quad \vec{\nabla}^2 \phi = \rho$$

$$\phi \sim Q \int d^3k \frac{e^{i\vec{k}\cdot\vec{x}}}{k^2} \sim \frac{1}{r} \quad r = |\vec{x}| \quad (\text{by dim analysis})$$

2) Newtonian potential same with $Q \rightarrow GM^2$

n extra dim : $\int d^3k \rightarrow \int d^{3+n}k \Rightarrow \frac{1}{r} \rightarrow \frac{1}{r^{1+n}}$

3d $\rightarrow$ (3+n)d due to KK excitations

massive particle:

$$\int d^3k \frac{e^{i\vec{k}\cdot\vec{x}}}{k^2 + m^2} \approx \int dL \int_{-1}^1 \cos\theta \frac{e^{ikr \cos\theta}}{k^2 + m^2} \sim \frac{1}{r} \int_0^\infty dk \frac{k \sin kr}{k^2 + m^2}$$

$\sim e^{-mr} \quad m > 0$

$\Rightarrow$ Yukawa potential : $\frac{e^{-mr}}{r}$ $m = 0, \frac{1}{R}, \frac{2}{R}, \dots$

multiplicity 2

$$\Rightarrow \frac{1}{r} \left\{ 1 + 2e^{-r/R} + 2e^{-2r/R} + \dots \right\} \equiv V$$

$R \rightarrow 0 \Rightarrow V \rightarrow \frac{1}{r} \quad (3d)$

$R \rightarrow \infty \Rightarrow V \sim \frac{1}{r} \left\{ 1 + 2 \left(\frac{1}{1 - e^{-r/R}} - 1 \right) \right\} = \frac{1}{r} \left\{ 1 + 2 \frac{e^{-r/R}}{1 - e^{-r/R}} \right\}$

$R \rightarrow \infty \Rightarrow V \sim \frac{2R}{r^2} \quad Q \rightarrow Q_5 = 2RQ_4$

**Generalizations, orbifolds,
symmetry breaking by compactification**

- Generalization

Non-abelian vectors $A_M^a \leftarrow$ gauge group

$$A_M^a = \begin{pmatrix} A_M^a \\ \phi^a \equiv A_4^a \end{pmatrix} \Rightarrow \text{Tr } F_{MN}^2$$

adjoint scalar

$$\int dy \text{Tr } F_{MN}^2 = \text{Tr } F_{\mu\nu}^2(x) + \text{Tr}(\phi)^2 + \sum_{n \neq 0} \left(\text{massive vectors} \right)$$

$$F_{\mu\nu}^n = \partial_\mu A_\nu^n - \partial_\nu A_\mu^n + \sum_m [A_\mu^m, A_\nu^{n-m}] \quad \text{Tr}(F_{\mu\nu}^n)^2$$

- More extra dims $n > 1$ KK excitations (besides the 0-modes)

$$A_M \Rightarrow \begin{pmatrix} A_M^\mu \\ \phi^i \end{pmatrix} \rightarrow \bullet \text{ massive vectors} \\ \rightarrow n-1 \text{ massive scalars (adjoints)}$$

$$G_{MN} = \begin{pmatrix} G_{\mu\nu} & G_{\mu i} \\ G_{\nu j} & G_{ij} \end{pmatrix}$$

0-modes: 4d graviton, n vectors, $\frac{n(n+1)}{2}$ scalars

KK-modes: spin-2 $G_{\mu\nu}^n$, $n-1$ vectors, $\frac{n(n-1)}{2}$ scalars

" $\frac{n(n+1)}{2} - 1 - (n-1)$

massive degrees of freedom: $5 + 3(n-1) + \frac{n(n-1)}{2} = 3n+2 + \frac{n(n-1)}{2} = \frac{n^2 + 5n + 4}{2}$

$$= \frac{(n+4)(n+1)}{2} = \frac{D(D-3)}{2}$$

$$\frac{D(D+1)}{2} - D - D = \frac{D^2 - 3D}{2}$$

gauge cond gauge law

~~graviton~~ graviton in D dim

Normalized fields: $\tilde{\Phi}_n = \Phi_n \sqrt{\frac{V}{V_{\text{int}}}} \leftarrow$ internal volume

Scalars G_{ij} : non-trivial compactification space
sizes and ^{shape} (complex structure) of the manifold

Ex. $n=2$ $\begin{pmatrix} G_{44} & G_{45} \\ G_{45} & G_{55} \end{pmatrix} \Rightarrow \begin{cases} \sqrt{G} : \text{radius} \\ G_{44}/G_{55}, G_{45}/G_{55} : \text{complex structure / shape} \end{cases}$

- Chirality and orbifolds

4d chirality from extra dims : non trivial

Higher dim spinor $\Rightarrow$ L,R symmetric in 4d

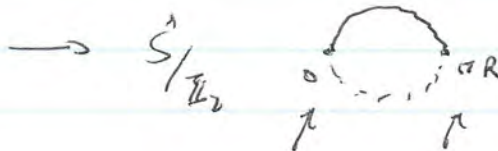
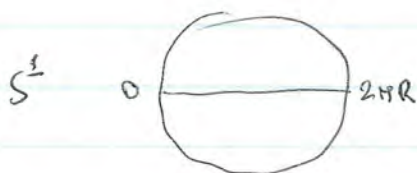
Ex. $D=5 \Rightarrow \gamma$ -matrices : $\{\gamma^M, \gamma^N\} = \gamma^{MN}$

$\Rightarrow$ 5d spinor $\equiv$ 4d Dirac spinor (L+R component)

Need non-trivial manifold

Orbifolds : special limits allowing chirality
at the expense of singular (fixed) points

Ex: one dim on an interval $\equiv S^1/\mathbb{Z}_2 \xrightarrow{y \equiv y+2\pi R}$
 $y \rightarrow -y$



Fixed points under the $\mathbb{Z}_2$

Boundary conditions

$$\Phi(y+2\pi R) \equiv \Phi(y) \Rightarrow \Phi(-y) = \begin{matrix} \text{even } (e) \\ \pm \Phi(y) \\ \text{odd } (o) \end{matrix}$$

$$\begin{aligned} \phi_{\text{even}} &= \sum_n \cos \frac{n}{R} y \phi_n \\ \phi_{\text{odd}} &= \sum_n \sin \frac{n}{R} y \phi_n \end{aligned} \quad \left\{ \begin{array}{l} \frac{|n|+|-n|}{R} \text{ KK} \\ \text{invariant under } y \rightarrow -y \text{ and } y \rightarrow y+2\pi R \\ \frac{|n|-|-n|}{R} \text{ KK} \end{array} \right.$$

ϕ_{odd} : no zero-modes $\phi_{\text{odd}}(0) = \phi_{\text{odd}}(\pi R) = 0$, $\phi_0 = 0$ $n: \text{odd}$

ϕ_{even} : $n: \text{even}$ $\phi_{\text{even}}(0) = \cancel{\phi_{\text{even}}(\pi R)} \phi_0 = -\phi_{\text{even}}(\pi R)$

Fermions

$$\begin{aligned} \psi(-y) &= \gamma^4 \psi(y) = \gamma^5 \psi(y) \Rightarrow \psi_L = \frac{1+\gamma^5}{2} \psi \text{ invariant} \\ \psi_R &= \frac{1-\gamma^5}{2} \psi \rightarrow -\psi_R \end{aligned}$$

$\Rightarrow$ 0-modes only for $\psi_L \Rightarrow$ 4d chirality

Gauge fields

$$A_M(-y) = \begin{pmatrix} A_\mu(y) \\ -A_4(y) \end{pmatrix} \Rightarrow \begin{matrix} A_\mu: \text{even} \\ A_4: \text{odd} \end{matrix}$$

$\Rightarrow A_\mu^{(n)}$ $n: \text{even}$ massive + 0-mode, no scalar

Metric

$$G_{MN} = \begin{pmatrix} G_{\mu\nu} & G_{\mu 4} \\ G_{4\nu} & G_{44} \end{pmatrix} \Rightarrow \begin{matrix} G_{\mu\nu}^{(n)} & n: \text{even} \\ G_{44} & \text{radion} \end{matrix}$$

$\begin{matrix} \text{odd} \\ \text{even} \end{matrix}$

no vectors

• fixed points: boundary states (twisted)

fixed points: singular (breakdown of EOM)

4d anomalies
 $\Rightarrow$ in general ~~extra~~ ^{localized} states ~~(sing)~~ at the fixed points (boundaries)
 model dependent (twisted in string theory)

$\Rightarrow$ notion of "branes"

• Coordinate dependent compactification, mass generation and symmetry breaking
 • Scherk-Schwarz eq

1) symmetry of higher dim theory $\Rightarrow$ general boundary conditions
 periodic up to symmetry transformation

e.g. U(1) phase: $\Phi(y+2\pi R) = e^{2i\pi\omega} \Phi(y)$ $\leftarrow$ complex field

$$\Rightarrow \Phi(y) = \sum_n e^{iy \frac{n+\omega}{R}} \phi_n$$

$$S_{1/2} \Rightarrow \phi_+ = \sum_n \cos y \frac{n+\omega}{R} \phi_n^+ \quad \phi_- = \sum_n \sin y \frac{n+\omega}{R} \phi_n^-$$

$\Rightarrow$ 0-mode gets a mass ω/R , $|n\rangle, | -n\rangle$ split (for ω arbitrary)

2) gauge symmetry breaking e.g. $SU(2) \rightarrow U(1)$ (τ_3 generator)

$$\bar{\Phi}(y+2\pi R) = e^{-2i\pi\omega t_3} \bar{\Phi}(y) e^{2i\pi\omega t_3} \quad \bar{\Phi} \equiv A_m$$

$$\Rightarrow \phi_+ \tau_3 \phi_-$$

$$\{\sigma^{\pm}, t_3\} = 0 \Rightarrow t_3 = \frac{\sigma_3}{2}$$

$$\bar{\Phi}_{\pm}(y+2\pi R) = e^{-4i\pi\omega t_3} \bar{\Phi}_{\pm}(y), \quad \phi_3(y+2\pi R) = \phi_3(y)$$

$$= [\cos 2\pi\omega - i\sigma_3 \sin 2\pi\omega] \phi_{\pm}(y)$$

$$\sigma^{\pm} = \frac{1}{2}(\sigma^1 \pm i\sigma^2) = \frac{1}{2} \begin{pmatrix} 1 & 1 \\ 0 & 0 \end{pmatrix} \pm \frac{i}{2} \begin{pmatrix} 0 & -i \\ i & 0 \end{pmatrix}$$

$$\Rightarrow \sigma^+ = \begin{pmatrix} 1 & 1 \\ 0 & 0 \end{pmatrix} \quad \sigma^3 \sigma^+ = \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix} \begin{pmatrix} 1 & 1 \\ 0 & 0 \end{pmatrix} = \begin{pmatrix} 0 & 1 \\ 0 & 0 \end{pmatrix} = \sigma^+$$

$$\sigma^3 \sigma^- = \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix} \begin{pmatrix} 0 & 0 \\ 1 & 1 \end{pmatrix} = \begin{pmatrix} 0 & 0 \\ -1 & 0 \end{pmatrix} = -\sigma^-$$

$$\Rightarrow \phi_{\pm}(y+2\pi R) = e^{\mp 2i\pi\omega} \phi_{\pm}(y)$$

$$\Rightarrow \phi_{\pm} \text{ gets a mass } (n \pm \omega)^2$$

$$\Rightarrow A_{\mu,n}^{\pm} \text{ absorbs } A_{4,n}^{\pm}, \quad A_{\mu,0}^{\pm} \text{ massless}, \quad A_{4,0}^{\pm} \text{ massless}$$

$$SU(2) \rightarrow U(1) \text{ by averaging if } (A_4) \sim \omega \sigma^3$$

3) supersymmetry breaking

use for $U(1)$ an R-symmetry

fermions (gravitino, gauginos) transform $\Rightarrow$ SUSY is broken

frequencies shifted by ω : $n + \omega \equiv m_n$

usually R-symmetry discrete $\Rightarrow \omega$ is quantized

special case $\omega = \frac{1}{2} \Rightarrow$ fermions antiperiodic $\Rightarrow \frac{1}{2}$ -integer frequencies
as in finite temperature $y \rightarrow$ time

Branes in extra dimensions

- Brane: hypersurface of p -spatial dimensions (p -brane)
 embedded in a higher dimensional bulk ($D = 4+n > p+1$)

Field theory: topological defect

String theory: D-branes defined by the ends of open strings

branes can carry "charge" (necessary for stability)

point particle $\equiv$ 0-brane can be electric source for gauge field (1-form)

$$e_0 \int dx^\mu A_\mu$$

string $\equiv$ 1-brane: for 2-form gauge potential (2-index tensor)

$$e_1 \int dx^\mu dx^\nu A_{\mu\nu}$$

$\vdots$
 p -brane $\rightarrow$ $(p+1)$ -form gauge potential

$$e_p \int \underbrace{dx^{\mu_1} \dots dx^{\mu_{p+1}}}_{\substack{\text{brane world volume} \\ \text{coordinates}}} A^{(p+1)}_{\mu_1 \dots \mu_{p+1}}$$

$$\int_{\Sigma^{p+1}} \epsilon^{\alpha_1 \dots \alpha_{p+1}} \partial_{\alpha_1} x^{\mu_1} \dots \partial_{\alpha_{p+1}} x^{\mu_{p+1}} A^{(p+1)}_{\mu_1 \dots \mu_{p+1}} \quad (*)$$

$\nearrow$
 new notation

$\rightarrow$ after defined brans

$$\int_{\Sigma^{p+1}} \epsilon^{\alpha_1 \dots \alpha_{p+1}} \partial_{\alpha_1} x^{\mu_1} \dots \partial_{\alpha_{p+1}} x^{\mu_{p+1}} A^{(p+1)}_{\mu_1 \dots \mu_{p+1}}$$

eg D3-brane in 10d: $ds^2 = f^{-1/2} dx_{||}^2 + f^{1/2} (dr^2 + r^2 d\Omega_3^2)$
 $\rightarrow$ unit 3-sphere

$f = 1 + \frac{4\pi g^2 L^4}{r^6}$: harmonic function $\Delta_6 \frac{1}{r^4} = \delta^6(x)$

two brans: gravity attractive force
 A repulsive " (line cm) $\rightarrow$ the way curved

Brane analogy with an extremal charged black hole

Charged BH (Reissner-Nordstrom ~~solution~~ solution)

$$ds^2 = - \left(1 - \frac{r_s}{r} + \frac{r_Q^2}{r^2} \right) dt^2 + \left(1 - \frac{r_s}{r} + \frac{r_Q^2}{r^2} \right)^{-1} dr^2 + r^2 d\Omega_2^2$$

$\uparrow$
2-sphere of unit radius
($d\theta^2 + \sin^2\theta d\phi^2$)

$$r_s = 2GM$$

$$r_Q^2 = e^2 G$$

$$\Rightarrow \text{two horizons at } r = r_{\pm} = \frac{1}{2} \left(r_s \pm \sqrt{r_s^2 - 4r_Q^2} \right)$$

$$A_\mu = \left(-\frac{e}{r}, \vec{0} \right)$$

extremal case: $r_+ = r_- \Rightarrow r_s = 2r_Q \Rightarrow \boxed{GM^2 = e^2}$

$$\Rightarrow ds^2 = - \left(1 - \frac{r_Q}{r} \right)^2 dt^2 + \underbrace{\left(1 - \frac{r_Q}{r} \right)^{-2} dr^2 + r^2 d\Omega_2^2}_{f^2(R) [dR^2 + R^2 d\Omega_2^2]}$$

$$r = R f(R) \Rightarrow \frac{dr}{1 - r_Q/r} = f(R) dR = \frac{r}{R} dR \Rightarrow \frac{dr}{r - r_Q} = \frac{dR}{R}$$

$$\Rightarrow r = r_Q + R, \quad f(R) = \frac{r}{R} = 1 + \frac{r_Q}{R}, \quad 1 - \frac{r_Q}{r} = \frac{R}{r} = \frac{1}{f(R)}$$

$$\Rightarrow ds^2 = - \frac{1}{f^2(R)} dt^2 + f^2(R) (dR^2 + R^2 d\Omega_2^2)$$

with $f(R) = 1 + \frac{r_Q}{R}$ = harmonic function

$$A_3 f(R) = r_Q \delta(x)$$

- 16" -

$\Rightarrow$ extremal BH $\equiv$ 0-brane

$\hookrightarrow$ point particle with charge $e = \sqrt{GM^2}$

force between two:

$$\begin{array}{ccc} \nearrow & \frac{GM^2}{r} & - \frac{e^2}{r} = 0 \quad \text{for } e^2 = \sqrt{GM^2} \\ \text{gravity} & & \downarrow \\ \text{attractive} & & \text{em - repulsive} \end{array}$$

no force $\Rightarrow$ stability

branes with charge in the extremal limit:

generalization of this situation

* $G_X \equiv$ 3-brane world

X^M ; $M=0, 1, \dots, 3+n$: space-time coordinates of the bulk

x^μ ; $\mu=0, 1, 2, 3$ coordinates of the 3-brane (our spacetime)

$Y^M(x)$: positions of the 3-brane in the extra dims

$(4+n)$ -dim metric $G_{MN}(X)$

Bulk fields : $B(X)$

Brane-localized fields : $\Phi(x)$

Vacuum state: $G_{MN} = \eta_{MN}$ (flat space bulk)

$$Y^M = \delta^M_\mu x^\mu$$

~~we can~~

→ describe fluctuations around the vacuum state



$$S_{\text{bulk}} = - \int d^{4+n} X \sqrt{|G|} \left\{ M_x^{2+n} R^{(4+n)} + \Lambda + \text{other bulk fields} \right\}$$

Induced metric on the brane:

$$ds^2 = G_{MN} dY^M(x) dY^N(x) = \underbrace{G_{MN} \frac{\partial Y^M}{\partial x^\mu} \frac{\partial Y^N}{\partial x^\nu}}_{g_{\mu\nu} : \text{induced metric}} dx^\mu dx^\nu$$

~~the~~ Symmetries of the EFT Action : general parametrization of the bulk
+ " " of the brane coordinates

⇒ ex. $\partial_\mu Y^M$: vector under bulk coordinate transformations
" " brane

$g_{\mu\nu}$: tensor under brane transfs
scalar " bulk

etc.

$$S_{\text{brane}} = - \int d^4x \sqrt{|g|} \left\{ T^{\mu\nu} + R^{(4)} + \frac{1}{2} g^{\mu\nu} D_\mu \Phi D_\nu \Phi + \frac{1}{2} F_{\mu\nu}^2 + \dots \right\}$$

gauge choice of brane world transf. $x^\mu \rightarrow x'^\mu(x)$

$$4 \text{ conditions} \Rightarrow Y^\mu(x) = x^\mu, \quad Y^m(x) \quad m=4, 5, \dots, 3+n$$

physical degrees of freedom
brane positions in the n -extra dims.

$$G_{MN} = \gamma_{MN} + \frac{1}{M_*^{4+n/2}} h_{MN} \quad \leftarrow \text{higher dim graviton}$$

$$\text{Brane fluctuations: } g_{\mu\nu} = G_{MN} \partial_\mu Y^N \partial_\nu Y^M = \gamma_{\mu\nu} + \partial_\mu Y^m \partial_\nu Y_m + \dots$$

$$\Rightarrow \det g = -1 - \partial_\mu Y^m \partial^\mu Y_m + \dots$$

T : brane tension ($T^{1/4} \ll M_*$ to neglect back reaction)

$$\text{canonically normalized field: } Z^m = T^{1/4} Y^m$$

$$\text{For } Y=0 \quad g_{\mu\nu} = G_{\mu\nu}(x^\lambda, Z^m=0) \quad \leftarrow \text{brane position}$$

$$\Rightarrow \text{For SM localized on the 3-brane} \quad S_{\text{SM}} = \int d^4x \int_{\text{SM}} \sqrt{g}(g_{\mu\nu}, \Phi, \dots)$$

$\Rightarrow$ Coupling of the brane fields to graviton:

$$S_{\text{int}} = \int d^4x \quad T_{\text{SM}}^{\mu\nu} \frac{h_{\mu\nu}(x_\mu)}{M_*^{4+n/2}}$$

$$\text{KK excitations: } h_{\mu\nu}^{(n)}(x,y) = \sum_{\vec{n}} e^{i \vec{n} y / R} \frac{h_{\mu\nu}^{(n)}}{\sqrt{2\pi R}}$$

e.g. for 1 extra dim

$\nearrow$
in general $n \rightarrow \vec{n}$
 $R \rightarrow \vec{R}$ for toroidal (or more general)
 $\sqrt{2\pi R} \rightarrow \sqrt{V_{\text{int}}}$

$$\Rightarrow S_{\text{int}} = \int d^4x \quad T_{\mu\nu}^{\text{SM}} \sum_{\mathbf{k}} \frac{h_{\mu\nu}^{\mathbf{k}}}{\sqrt{V_n}} \quad \left(M_p^2 = M_*^{2+n} V_n \right)$$

4d M_p

int volume

$\Rightarrow$ each KK mode couples with the same strength M_p
(as the 0-mode)

$\rightarrow$ brane charge from p. 16

* Exchange of KK modes between brane sources



Ex of bulk gauge field or bulk graviton

propagator: $\frac{1}{k^2 + m_0^2 + \frac{n^2}{R^2}}$ tensor structure (one extra dim)

m_0 : 5d mass

k : 4d momentum

$$\sum_{n \neq 0} \frac{1}{k^2 + m_0^2 + \frac{n^2}{R^2}} = \frac{\pi}{a} \coth \pi R a \Rightarrow$$

$$\sum_{\text{KK-modes } \neq 0} = \left(\frac{\pi}{R a} \coth \pi R a - \frac{1}{R^2 a} \right) R^2 g^{\mu\nu} \quad a = \sqrt{k^2 + m_0^2}$$

4d coupling $\frac{1}{g}$

$$\Sigma = g^{\mu\nu} \left(\frac{\pi R}{a} \coth \pi R a - \frac{1}{a^2} \right) \quad (\text{can be } \frac{1}{M_p} \text{ for gravity})$$

$\coth x = \begin{cases} \pm 1 & x \rightarrow \pm \infty \\ \frac{1}{x} + \sum_{n=1}^{\infty} \frac{2^{2n} B_{2n} x^{2n-1}}{(2n)!} & x \rightarrow 0 \end{cases}$

Bernoulli numbers

$$= \frac{1}{x} + \frac{x}{3} + \dots$$

limits:

• 4d limit $R \rightarrow 0 \Rightarrow \Sigma \rightarrow 0$

• $k=0$ (0-momentum) $\Rightarrow a=m_0$

$m_0(k=0) \Rightarrow \Sigma \rightarrow \frac{g^2}{3} \pi^2 R^2$: low energy eff operator of the form $(\text{current})^2$
 $\nwarrow$ $T_{\mu\nu}$ for gravity

• $m_0=0 \Rightarrow a=k (= \sqrt{k^2})$

• $k \rightarrow \infty \Rightarrow \Sigma \sim \frac{g^2 \pi R}{k} \sim \frac{g_5^2}{2k}$ where $g^2 = \frac{g_5^2}{2\pi R}$

e.g. $\frac{1}{g_5^2} F^2 \rightarrow \frac{2\pi R}{g_5^2} = \frac{1}{g_4^2}$ or $M_P^2 = M_*^3 2\pi R$
 $\nearrow g_4^2$ $\nearrow g_5^2$ for gravity

more dims: $2\pi R \rightarrow V_n$: volume of n -extra dims

• More than 1 extra dim $n \geq 2$: sum over $\vec{n}$ diverges

for $n=2$ logarithmic; for $n \geq 3$ power Λ^{n-2}

$\Rightarrow$ need regularization: with S : $e^{-S\vec{n}^2}$